

# Phase relations of the iron carbides $\text{Fe}_2\text{C}$ , $\text{Fe}_3\text{C}$ , $\text{Fe}_7\text{C}_3$ at the Earth's core conditions

N.E. Sagatov<sup>1\*</sup>, P.N. Gavryushkin<sup>1</sup>, T.M. Inerbaev<sup>1</sup> and K.D. Litasov<sup>1</sup>

<sup>1</sup>*Sobolev Institute of Geology and Mineralogy, Siberian Branch of the Russian Academy of Sciences, Novosibirsk, 630090, Russia*

Keywords: iron carbides, quasiharmonic approximation, crystal structure prediction.

\*e-mail: sagatovnye@igm.nsc.ru, sagatino23@gmail.com

The Fe-C is one of the important systems in the study of the composition and structure of the Earth's core [1, 2]. A number of intermediate compounds were considered as a possible carbon-bearing phase in the core:  $\text{Fe}_3\text{C}$ ,  $\text{Fe}_3\text{C}_2$ ,  $\text{Fe}_5\text{C}_2$ ,  $\text{Fe}_7\text{C}_3$  и  $\text{Fe}_2\text{C}$  [3-6]. It was assumed that among these carbides  $\text{Fe}_2\text{C}$  is most stable under the Earth's inner core conditions [6-7], however this conclusion is based on calculations at 0 K.

In the present work we investigated the stability of various iron carbides at pressures and temperatures of the Earth's core based on the calculations within the density functional theory, quasiharmonic approximation and structure prediction methods.

The crystal structure predictions performed for  $\text{Fe}_2\text{C}$ ,  $\text{Fe}_3\text{C}$ , and  $\text{Fe}_7\text{C}_3$  revealed two new phases:  $\text{Fe}_3\text{C}$ - $C2/m$  and  $\text{Fe}_7\text{C}_3$ - $C2/m$  which are thermodynamically stable above 300 GPa at 0 K.

The energy relationships between the predicted crystal structures are shown in Fig. 1. According to the calculated convex hull,  $\text{Fe}_7\text{C}_3$ - $C2/m$  is not stable and decomposes into a mixture of  $\text{Fe}_3\text{C} + 2\text{Fe}_2\text{C}$  at 0 K.  $\text{Fe}_3\text{C}$ - $C2/m$  is stable relative to a mixture of  $\text{Fe}_2\text{C} + \text{Fe}$  at 0 K.

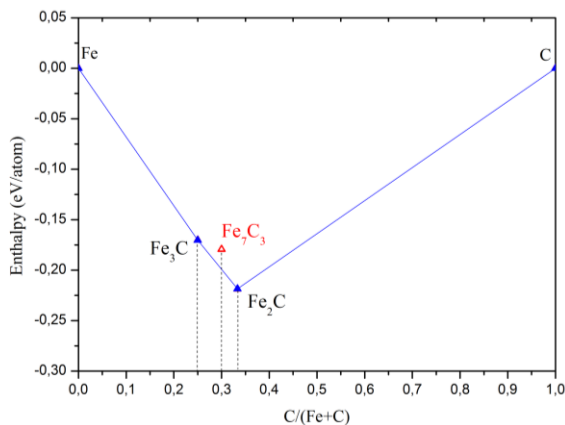


Figure 1. Convex hulls of the Fe-C system at 400 GPa

Using the quasiharmonic approximation, we investigated the effect of temperature on the stability of iron carbides (Fig. 2). According to the results, at P-T parameters of the Earth's inner core,  $\text{Fe}_3\text{C}$  is thermodynamically stable in the form of  $I-4$ ,  $\text{Fe}_7\text{C}_3$  – in the form of a newly predicted  $C2/m$  phase, and  $\text{Fe}_2\text{C}$  – in the form of  $Pnma$ -II, predicted by Bazhanova et al. [6]. Iron carbide  $\text{Fe}_7\text{C}_3$  is unstable relative to a mixture of  $\text{Fe}_3\text{C} + 2\text{Fe}_2\text{C}$  at high temperatures (Fig. 2). In this case, the stabilization of the reaction products is achieved due

to the transition from cementite to  $I-4$  phase. This transition also stabilizes  $\text{Fe}_3\text{C}$  relative to a mixture of  $\text{Fe} + \text{Fe}_2\text{C}$  over the entire range of pressures and temperatures. At 329 GPa, corresponding to the boundary between the inner and outer core, and 0 K, the difference in the Gibbs free energies of  $\text{Fe} + \text{Fe}_2\text{C}$  and  $\text{Fe}_3\text{C}$  is 53.7 meV/f.u. and this value increases with increasing temperature.

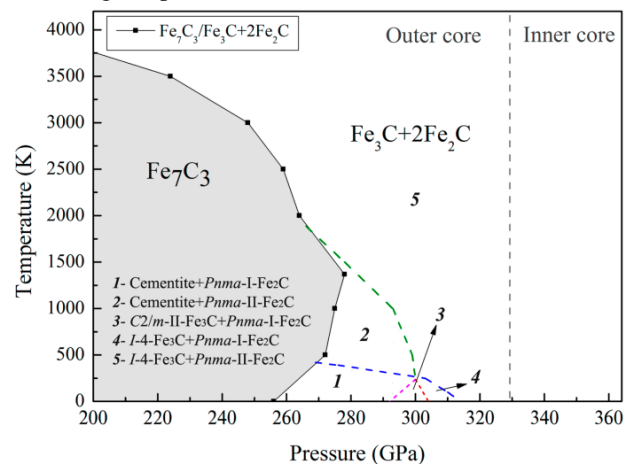


Figure 2. The boundary of the reaction  $\text{Fe}_7\text{C}_3 \rightarrow \text{Fe}_3\text{C} + 2\text{Fe}_2\text{C}$  in P-T coordinates. The numbers indicate the stability fields of the mixture  $\text{Fe}_3\text{C} + 2\text{Fe}_2\text{C}$  with different combinations of the structures  $\text{Fe}_3\text{C}$  and  $\text{Fe}_2\text{C}$ .

**Acknowledgments:** This work was supported by the Russian Science Foundation (project no. 17-17-01177).

- [1] B.J. Wood, *Earth Planet. Sci. Lett.*, 1993, **117**, 593.
- [2] B.J. Wood, J. Li, A. Shahar, *Rev. Mineral. Geochem.*, 2013, **75**, 231.
- [3] O. Lord, M. Walter, R. Dasgupta, D. Walker, S. Clark, *Earth Planet. Sci. Lett.*, 2009, **284**, 157.
- [4] Y. Nakajima, E. Takahashi, T. Suzuki, K. Funakoshi, *Phys. Earth Planet. Inter.*, 2009, **174**, 202.
- [5] G.L. Weerasinghe, R. Needs, C.J. Pickard, *Phys. Rev. B*, 2011, **84**, 174110.
- [6] Z. G. Bazhanova, A. R. Oganov and O. Gianola, *Phys.-Usp.*, 2012, **55**, 489.
- [7] A.R. Oganov, R.J. Hemley, R.M. Hazen and A.P. Jones, *Rev. Mineral. Geochem.*, 2013., **75**, 47.